

A composite space image featuring a view of Earth from space, showing a curved horizon and a layer of white clouds. In the upper left, a crescent moon is visible against the dark background of space. In the upper center, a bright sun is partially obscured by a reddish lens flare. The text 'ASTRO BIOLOGY & WEBB TRANSIT TECHNIQUES' is overlaid on the right side of the image in a yellow, outlined, sans-serif font. A thin horizontal line is positioned between the words '& WEBB' and 'TRANSIT'.

ASTRO
BIOLOGY
& WEBB
TRANSIT
TECHNIQUES



1
00:00:10,150 --> 00:00:08,390
james webb is going to characterize the

2
00:00:13,190 --> 00:00:10,160
atmospheres of exoplanets using a

3
00:00:15,190 --> 00:00:13,200
technique called transit spectroscopy if

4
00:00:17,349 --> 00:00:15,200
it's watching a planet transit across

5
00:00:19,029 --> 00:00:17,359
its star so the planet passes directly

6
00:00:21,349 --> 00:00:19,039
in front of its star from our point of

7
00:00:23,990 --> 00:00:21,359
view a really sensitive telescope can

8
00:00:26,790 --> 00:00:24,000
actually pick up those tiny tiny dips in

9
00:00:28,870 --> 00:00:26,800
the starlight and measure that a planet

10
00:00:31,189 --> 00:00:28,880
is there that's how we discovered a lot

11
00:00:33,670 --> 00:00:31,199
of the planets that we found planets

12
00:00:36,229 --> 00:00:33,680
have atmospheres and atmospheres can

13
00:00:37,990 --> 00:00:36,239

actually let some of the light pass

14

00:00:39,350 --> 00:00:38,000

through them on the way to our

15

00:00:40,869 --> 00:00:39,360

telescopes

16

00:00:42,709 --> 00:00:40,879

there are different gases in these

17

00:00:45,830 --> 00:00:42,719

atmospheres that can absorb different

18

00:00:47,750 --> 00:00:45,840

colors of light and so by using a really

19

00:00:50,150 --> 00:00:47,760

sensitive telescope that can measure

20

00:00:52,470 --> 00:00:50,160

what colors of light get absorbed when

21

00:00:53,590 --> 00:00:52,480

the planet is being backlit by its star

22

00:00:55,590 --> 00:00:53,600

and some of that light is passing

23

00:00:57,830 --> 00:00:55,600

through the atmospheres then we can

24

00:00:59,590 --> 00:00:57,840

learn what's in those atmospheres what

25

00:01:01,349 --> 00:00:59,600

gases are there that's absorbing those

26

00:01:02,389 --> 00:01:01,359

various colors of light what are they

27

00:01:04,070 --> 00:01:02,399

made of

28

00:01:06,310 --> 00:01:04,080

there's another technique called the

29

00:01:08,469 --> 00:01:06,320

phase curve technique the phase curve

30

00:01:10,789 --> 00:01:08,479

technique looks at the planet and the

31

00:01:12,870 --> 00:01:10,799

star when the planet is not in front of

32

00:01:14,149 --> 00:01:12,880

the star you got the signal from both

33

00:01:16,710 --> 00:01:14,159

and it also looks at them when the

34

00:01:18,310 --> 00:01:16,720

planet passes directly behind the star

35

00:01:20,469 --> 00:01:18,320

so in that case you can only see the

36

00:01:22,310 --> 00:01:20,479

signal from the star and by taking the

37

00:01:23,510 --> 00:01:22,320

difference of those two signals you can

38

00:01:25,910 --> 00:01:23,520

figure out what the signal from the

39

00:01:28,070 --> 00:01:25,920

planet was if you have a sufficiently

40

00:01:29,670 --> 00:01:28,080

sensitive instrument and from that

41

00:01:32,149 --> 00:01:29,680

technique we can also learn things about

42

00:01:33,510 --> 00:01:32,159

what the planet is made out of and also

43

00:01:35,590 --> 00:01:33,520

some interesting things about the

44

00:01:38,390 --> 00:01:35,600

planet's temperature and all this

45

00:01:40,469 --> 00:01:38,400

together tells us a lot of information

46

00:01:43,650 --> 00:01:40,479

about what these planets are like and